

# Gaussian Processes for Systematic Mitigation

## A Bayesian Method to Mitigate the Effects of Unmodelled Time-Varying Systematics for Global 21-cm Cosmology Experiments

**Christian J. Kirkham**<sup>1,2</sup>, Dominic J. Anstey<sup>1,2</sup>, Eloy de Lera Acedo<sup>1,2</sup>

<sup>1</sup>Cavendish Astrophysics, University of Cambridge

<sup>2</sup>Kavli Institute for Cosmology, University of Cambridge

URSI AT-RASC, 23 May 2024

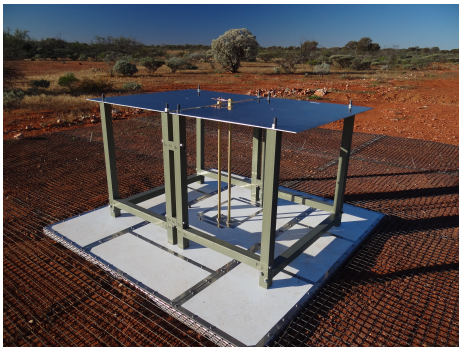


# Table of Contents

- 1 Introduction
- 2 Systematics in the REACH System
- 3 Gaussian Processes
- 4 Results
- 5 Further Work



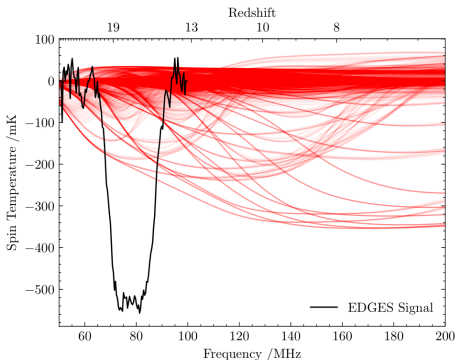
# EDGES (Experiment to Detect the Global EoR Signature)



- EDGES is a low frequency radio experiment to detect the global sky-averaged 21-cm signal
- Collaboration between ASU and MIT
- Located in the Murchison Radio-astronomy Observatory in Western Australia



# EDGES (Experiment to Detect the Global EoR Signature)



- Claimed detection of 21-cm signal at 78 MHz (Bowman+18)
- Unusually deep signal, requiring exotic physics
- Concerns that there is a residual systematic in the data (Hills+18, Sims and Pober 2020)



# REACH (Radio Experiment for the Analysis of Cosmic Hydrogen)

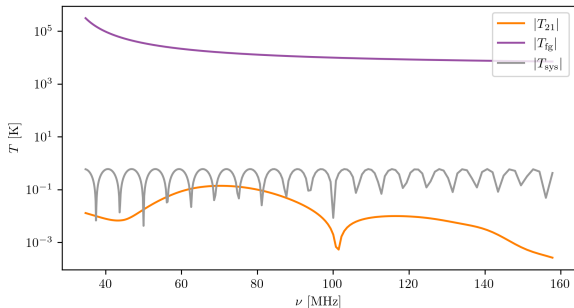


Credit: Saurabh Pegwal

- REACH is a global signal experiment designed to verify the EDGES detection
- Cambridge University, Stellenbosch University and others
- Located in the Karoo Desert on the SARAO facility in South Africa



# Global Experiment Challenges



Credit: Harry Bevins

- **Galactic foregrounds**  
Smooth synchrotron emission  
 $\sim 10^5$  K
- **Instrumental systematics**  
Generally sinusoidal  
 $\sim 10^0$  K
- **21-cm Signal**  
Not spectrally smooth  
 $\sim 10^{-2}$  K



# Time Varying Systematics in the REACH System



- Some systematics are expected to be static, others vary with time
  - Reflections from the soil vary with rainfall (Bevins+21, 22)
  - Impedences in the system are temperature dependent
  - Improperly modelled beams can cause systematics which vary with the galactic foreground



# Modelling Time Varying Systematics

- Want to simulate a systematic to test robustness of data analysis techniques
- Model systematic as a sinusoid whose amplitude is modulated by the power of the galactic foreground

$$T_{\text{sys}}(\nu) = A_{\text{sys}} \left( \frac{\nu}{\nu_{0,\text{sys}}} \right)^{-\alpha_{\text{sys}}} \sin \left( \frac{2\pi\nu}{P_{\text{sys}}} + \phi_{\text{sys}} \right) \quad (1)$$

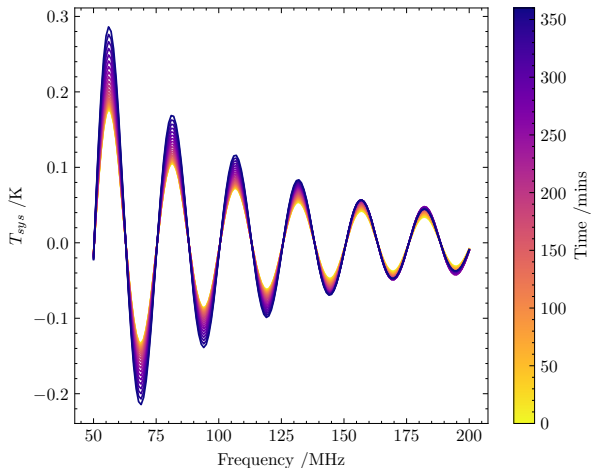
- Systematic amplitude is varied as

$$A_{\text{sys}}(t_j) = A_{\text{sys}}(t_0) \cdot \frac{T_{\text{fg}}(\nu = \nu_0, t = t_j)}{T_{\text{fg}}(\nu = \nu_0, t = t_0)} \quad (2)$$





# Modelling Time Varying Systematics



# Standard REACH Pipeline

- Fully Bayesian forward model of antenna beam, galactic foregrounds and 21-cm signal (Anstey+21, 22)
- Uses a Gaussian likelihood with uncorrelated noise,  $\sigma_0$

$$\log \mathcal{L} = \sum_i -\frac{1}{2} \log(2\pi\sigma_0^2) - \frac{1}{2} \left( \frac{T_{\text{data}}(\nu_i) - (T_{\text{fg}}(\nu_i) + T_{21}(\nu_i) + T_{\text{CMB}})}{\sigma_0} \right)^2 \quad (3)$$

- Will be referred to as the “Standard Pipeline”



# Gaussian Processes

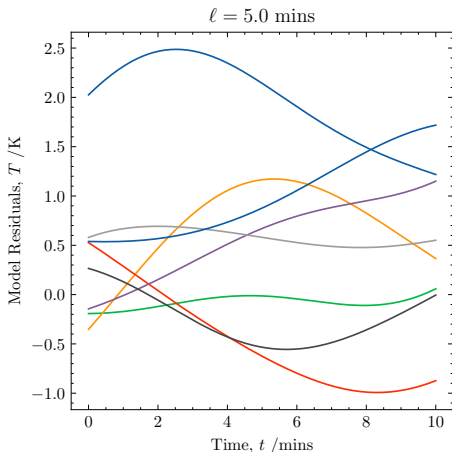
- Define a GP as a collection of random variables which have consistent joint Gaussian distributions (Rasmussen 2004)
- Distribution is defined by the covariance matrix, or covariance function

$$\mathbf{C}_{ij} = K(t_i, t_j) \quad (4)$$

- $K$  is called the “kernel” – form can be chosen freely



# Squared Exponential Kernel



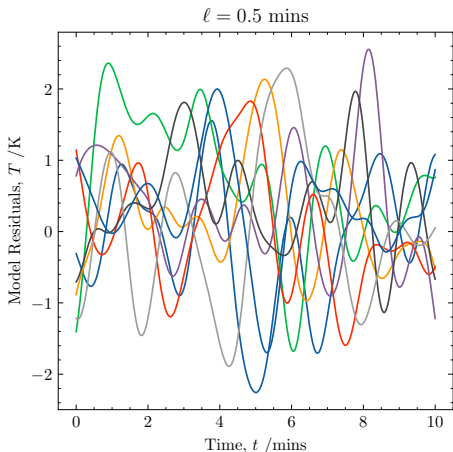
- In this work I use a squared exponential kernel with a white noise term

$$\mathbf{C}_{ij} = \sigma_{0,\text{GP}}^2 + \sigma_{\text{SE}}^2 \exp\left(-\frac{|t_i - t_j|^2}{2\ell^2}\right) \quad (5)$$

- Uncorrelated noise,  $\sigma_{0,\text{GP}}$ , scale factor of the kernel,  $\sigma_{\text{SE}}$ , and the covariance length,  $\ell$
- Define smooth functions



# Squared Exponential Kernel



- In this work I use a squared exponential kernel with a white noise term

$$\mathbf{C}_{ij} = \sigma_{0,\text{GP}}^2 + \sigma_{\text{SE}}^2 \exp\left(-\frac{|t_i - t_j|^2}{2\ell^2}\right) \quad (5)$$

- Uncorrelated noise,  $\sigma_{0,\text{GP}}$ , scale factor of the kernel,  $\sigma_{\text{SE}}$ , and the covariance length,  $\ell$
- Define smooth functions



# Gaussian Processes in the REACH Pipeline

- Hereafter called the “Gaussian Process pipeline” (Kirkham+24)

$$\mathcal{L}(\theta) = \frac{1}{\sqrt{(2\pi)^n |\mathbf{C}|}} \exp\left(-\frac{1}{2}(\mathbf{D} - \mathbf{M}(\theta))^T \mathbf{C}^{-1}(\mathbf{D} - \mathbf{M}(\theta))\right) \quad (6)$$

- $\mathbf{D}$  is the data vector,  $\mathbf{M}$  is the model (foreground, beam model, signal model)
- Gaussian process covariance matrix,  $\mathbf{C}$ , fits covariance of model residuals between time bins



# Simulating Data

- We model the global signal as a Gaussian

$$T_{\text{sg}}(\nu) = -A_{21} \exp\left(-\frac{(\nu - \nu_{0,21})^2}{2\sigma_{21}^2}\right) \quad (7)$$

- Simulated data produced with “true” signal parameters  $A_{21} = 0.155$  K,  $\sigma_{21} = 15$  MHz and  $\nu_{0,21} = 85$  MHz
- 21-cm signal is fitted for using same Gaussian model, with three free parameters:  $A_{21}$ ,  $\sigma_{21}$  and  $\nu_{0,21}$



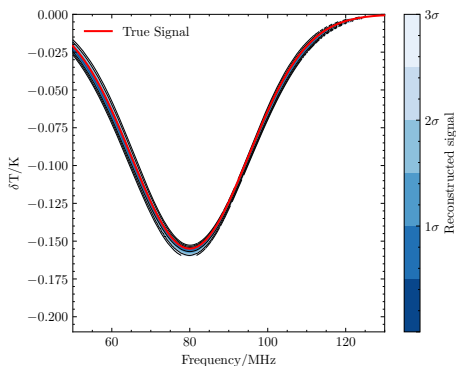
# Varying Systematic parameters

- Will test both pipelines with a range of simulated systematics
- Data generated for 24 time bins of length 15 minutes – 6 hours of observation in total
- Posterior distribution was then sampled using POLYCHORD

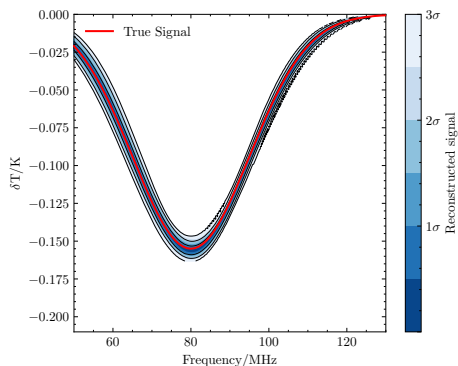




# Comparison of Signal Posteriors - No Systematic



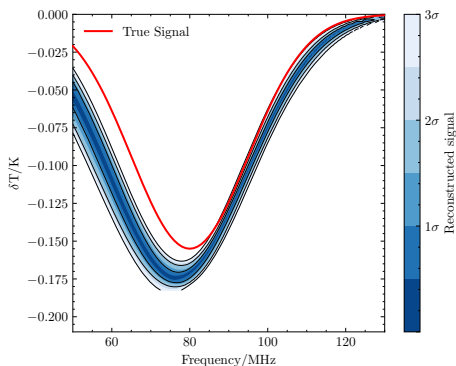
Standard pipeline



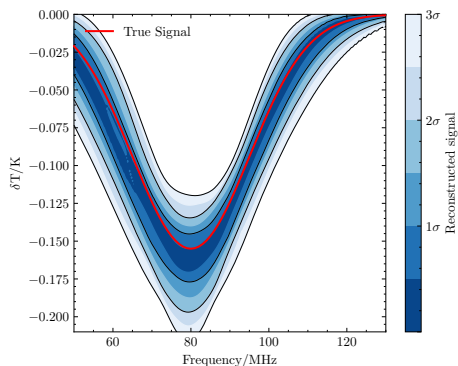
Gaussian Process pipeline



# Comparison of Signal Posteriors - With Systematic



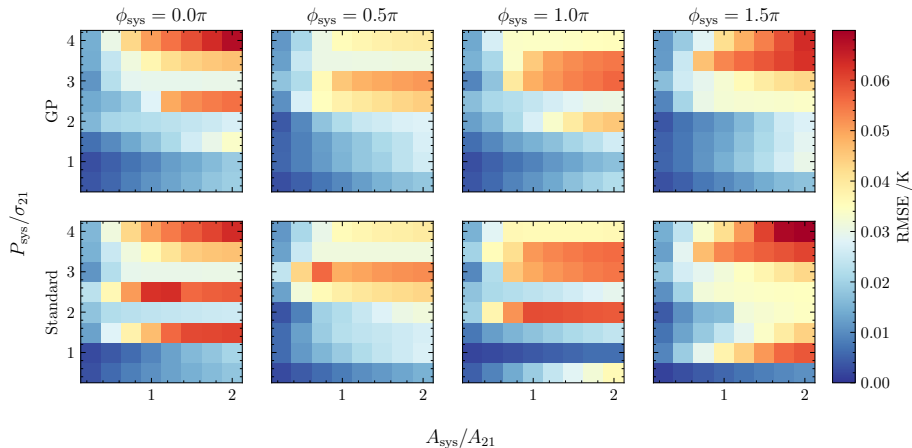
Standard pipeline



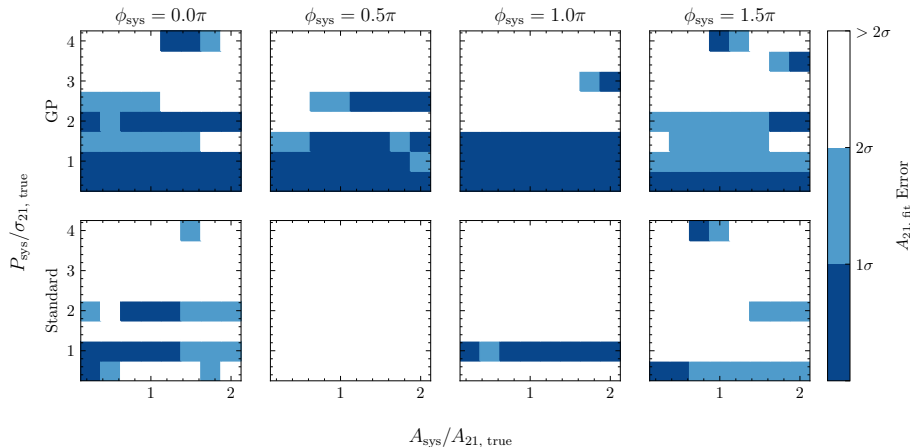
Gaussian Process pipeline



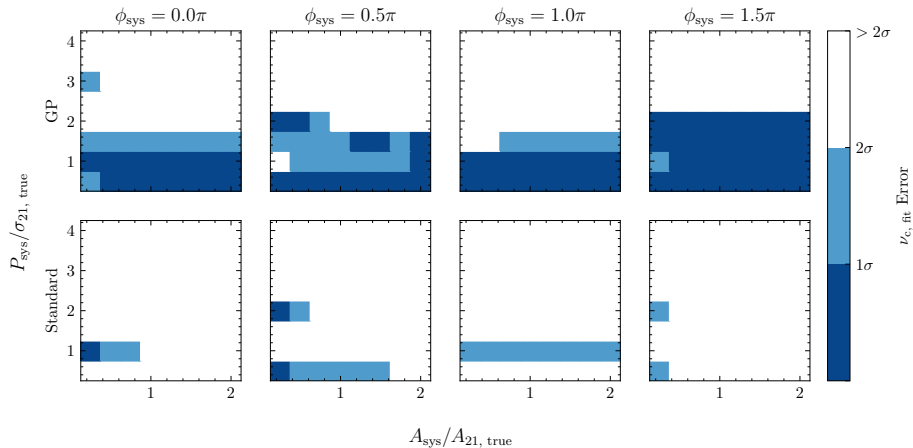
## RMSE



# Fitted Signal Amplitude error



# Fitted Centre Frequency error



# Gaussian Process Regression

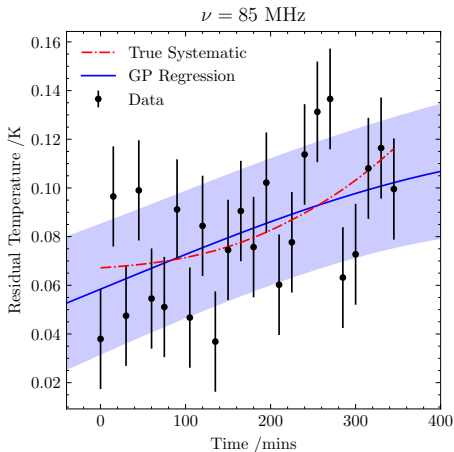
- Can also use Gaussian Processes for regression
- Mean function of the Gaussian Process posterior given by

$$\mu(\mathbf{t}_{\text{pred}}) = K(\mathbf{t}_{\text{pred}}, \mathbf{t}_{\text{data}})K(\mathbf{t}_{\text{data}}, \mathbf{t}_{\text{data}})^{-1}\mathbf{T}_{\text{data}}, \quad (8)$$

- Can be used to predict temperature,  $\mathbf{T}_{\text{pred}}$ , of model residuals at some time,  $\mathbf{t}_{\text{pred}}$



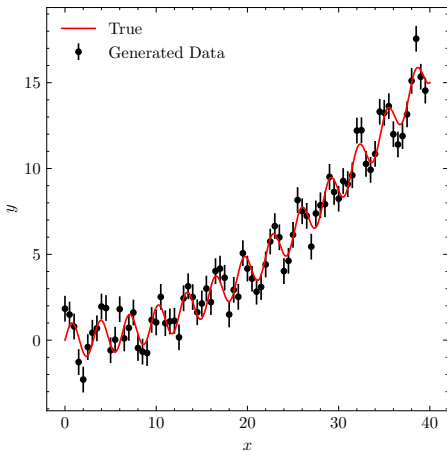
# Gaussian Process Regression



- Use weighted mean Gaussian Process hyperparameters to get a smooth fit to model residuals
- Could possibly be used to inform future time-varying systematic models



# Automatic Kernel Selection

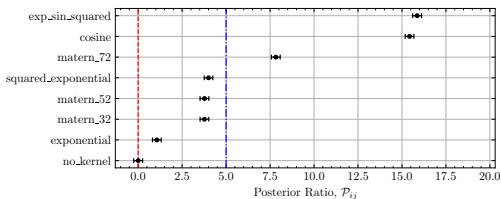


- Kernel choice is arbitrary so can use Bayes factor to inform us (Hee+15, Kroupa+24)
- Basic test with a quadratic curve with a sinusoidal residual
- Fit for quadratic but not sinusoid





# Automatic Kernel Selection



- Uses PolyChord to sample over all kernels using a choice parameter,  $c$
- Uses the posterior ratio of  $c$  as a proxy for Bayes Factor

$$\mathcal{P}_{ij} = \log \frac{\Pr(c = j | \mathbf{D}, \mathbf{M})}{\Pr(c = i | \mathbf{D}, \mathbf{M})} \quad (9)$$



# Summary

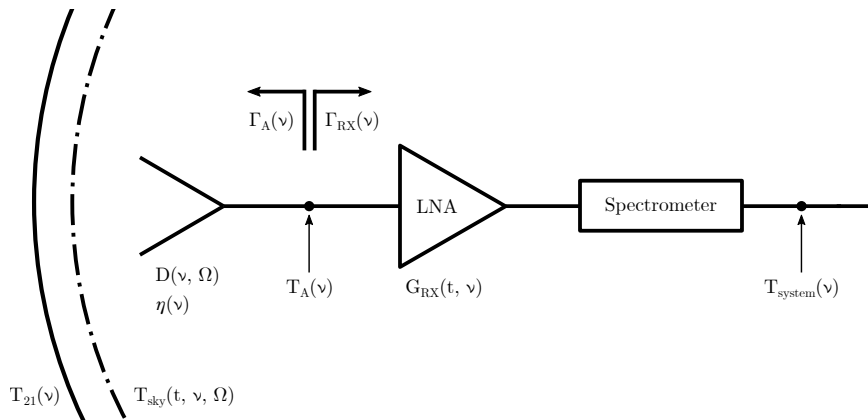
- Using Gaussian Processes to account for time correlated residuals improves fitting
- General method – no systematic model required
- Regression can inform future models of systematics
- Automatic Kernel Selection can use the data to select the most appropriate kernel



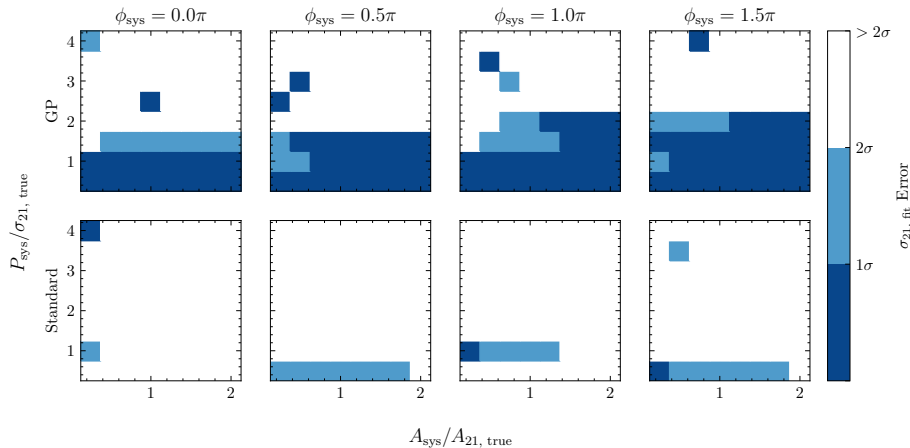
For more information



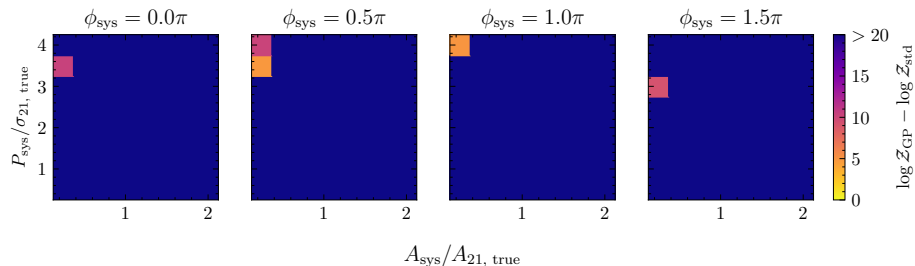
# Antenna and Receiver System



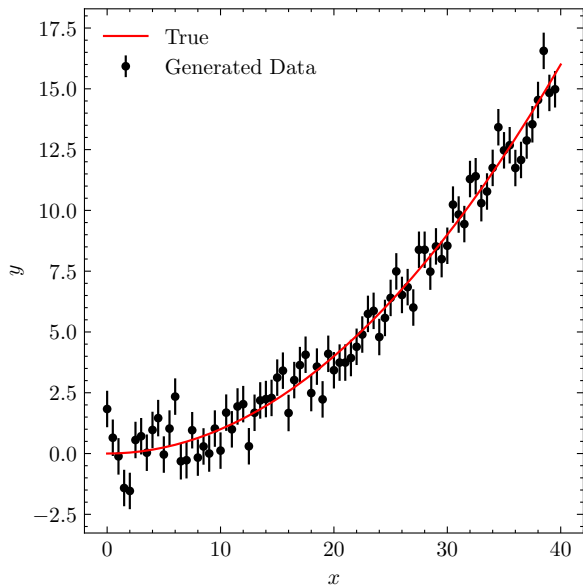
# Signal Width



# Bayes Factor



# Automatic Kernel Selection - No Residual



# Automatic Kernel Selection - No Residual

